

We have derived the formula for the photon scattering rate in the case of weak narrowband excitation,

$$R_{sc} = \frac{\omega_R^2}{\Gamma} \frac{|\mathbf{r}/c|^4}{|\mathbf{r}/c|^2 + \Delta^2}$$

in terms of the Rabi frequency $\omega_R = \sqrt{n+1} \omega_1$,

and the n-photon Rabi frequency $\omega_1 = -\frac{2}{\hbar} \mathbf{d} \cdot \mathbf{e} \sqrt{\frac{\hbar \omega_{eg}}{2 \epsilon_0 V}}$

To express this in terms of classical-field properties,

we note that (using $(\mathbf{d} \cdot \mathbf{e})^2 \rightarrow d_x^2 = \frac{1}{3} d^2$),

$$\omega_R^2 = (\mathbf{d} \cdot \mathbf{e})^2 \frac{2 \omega_{eg} \hbar (n+1)}{\epsilon_0 \hbar V} = \frac{d^2 \omega_{eg}^2}{3 \pi \epsilon_0 \hbar c^3} \frac{2 \pi c^3}{V \omega_{eg}^2} (n+1) = \Gamma^2 \frac{6 \pi c^3}{3 V \omega_{eg}^2} (n+1)$$

where $A = \Gamma = \frac{g_0}{g_c} \frac{d^2 \omega_{eg}^2}{3 \pi \epsilon_0 \hbar c^3} = \frac{d^2 \omega_{eg}^2}{9 \pi \epsilon_0 \hbar c^3}$ for an S-P transition

For large n , $n+1 \approx n$, $V = \text{area} \cdot \text{length}$, $\text{time} = \frac{\text{length}}{c}$

$$I = \frac{\text{energy}}{\text{area} \cdot \text{time}} = \frac{\hbar \omega n}{V/c} \approx \frac{\hbar c \omega (n+1)}{V} \approx \frac{\hbar c \omega_{eg} (n+1)}{V}$$

$$\frac{\omega_R^2}{\Gamma^2} = I \frac{6 \pi c^2}{\hbar \omega_{eg}^3 \Gamma}$$

The last term has units of $(\text{intensity})^{-1}$, we define

$$I_s = \frac{\hbar \omega_{eg}^3 \Gamma}{12 \pi c^2} \quad \text{saturation intensity}$$

This allows us to express the Rabi frequency as

$$\omega_R^2 = \frac{I}{2I_s} \Gamma^2$$

and the scattering rate becomes

$$R_{sc} = \frac{\omega_R^2}{\Gamma} \frac{(\Gamma/2)^2}{(\Gamma/2)^2 + \Delta^2} = \frac{I}{I_s} \frac{(\Gamma/2)^2}{(\Gamma/2)^2 + \Delta^2} \frac{\Gamma}{2}$$

We have derived the formula for $\omega_R^2 \ll \Gamma^2$,
i.e. for $I \ll I_s$. It remains valid for

$$p := \frac{I}{I_s} \frac{(\Gamma/2)^2}{(\Gamma/2)^2 + \Delta^2} \ll 1$$

p is called the saturation parameter

The scattering cross section is the inverse ratio
intensity and scattered power. On resonance we have

$$\sigma_0 = \frac{h \nu_{eg} R_{sc}}{I} = \frac{h \nu_{eg} \Gamma}{2 I_s} = \frac{6\pi c^2}{\omega_{eg}^2} = \frac{6\pi}{k^2} = 6\pi \lambda^2 = \frac{3}{2\pi} \lambda^2$$

$$\sigma_0 = 6\pi \lambda^2 = \frac{3}{2\pi} \lambda^2 \quad \text{resonant cross section}$$

($\lambda = \frac{\lambda}{2\pi}$). Note that the resonant cross section

is a) independent of linewidth and b) much larger
than the atomic size.

The off-resonant scattering cross section is

$$\sigma(\Delta) = \sigma_0 \frac{(\Gamma/2)^2}{(\Gamma/2)^2 + \Delta^2}$$

The saturation intensity can be rewritten as

$$I_s = \frac{\hbar \omega_0 \Gamma/2}{\frac{3}{2\pi} \lambda^2} = \frac{\hbar \omega_0 \Gamma/2}{\sigma_0}$$